

**DEBRIS: Disc Emission via a Bias-free Reconnaissance in the  
Infrared/Sub-millimetre**

**Principal Investigator:** Brenda Matthews

**Institution:** Herzberg Institute of Astrophysics

**E-mail address:** [brenda.matthews@nrc-cnrc.gc.ca](mailto:brenda.matthews@nrc-cnrc.gc.ca)

**Co-Investigators:** Jane Greaves, St. Andrews University, UK

Pierre Bastien, Universite Montreal, Canada

Andy Biggs, University of Edinburgh, UK

Jerome Bouvier, Laboratoire d'Astrophysique de  
Grenoble, France

Harold Butner, James Madison University, USA

Bill Dent, UK Astronomy Technology Centre, UK

James Di Francesco, National Research Council of  
Canada

Gaspard Duchene, University of California, Berkeley,  
USA

Jochen Eisloffel, Thuringer Landessternwarte,  
Tautenburg, Germany

James Graham, University of California, Berkeley, USA

Paul Harvey, University of Texas at Austin, USA

Peter Hauschildt, Hamburger Sternwarte, Germany

Wayne Holland, UK Astronomy Technology Centre, UK

Jonti Horner, The Open University, UK

Eduardo Ibar, University of Edinburgh, UK

Rob Ivison, UK Astronomy Technology Centre and  
Institute for Astronomy, University of Edinburgh,  
UK

Doug Johnstone, National Research Council of Canada

Paul Kalas, University of California, Berkeley, USA

Jean-Francois Lestrade, Observatoire de Paris, France

Amaya Moro–Martin, Department of Molecular and  
Infrared Astrophysics, CSIC, Spain

Neil Phillips, University of Edinburgh, UK

Bruce Sibthorpe, Cardiff University, UK

Stephane Udry, Geneva Observatory, Switzerland

Paul van der Werf, Leiden Observatory, University of  
Leiden, The Netherlands

David Wilner, Smithsonian Astrophysical Observatory,  
USA

Mark Wyatt, University of Cambridge, UK

Ben Zuckerman, University of California, Los Angeles,  
USA

**Science Category:** Galactic: Circumstellar/Debris disks

**Observing Modes:** Pacs Photometer, Spire Photometer

**Hours Requested:** 353.9

**Abstract:**

Debris discs are belts of dust particles created from collisions of planetesimals (comets and asteroids) in extrasolar planetary systems. The prototype disc around Vega was discovered ~25 years ago by IRAS, and ~20 discs have been imaged to date, primarily by HST, SCUBA and Spitzer. Despite the relatively low numbers, debris discs are seen to be extraordinarily diverse in character, including systems with vast populations of comets or with perturbations by planets at tens of AU from the host star. Due to this low number, however, our knowledge of debris discs is incomplete; there has been no unbiased survey specifically designed to image a large number of discs. Spitzer greatly improved our understanding of the disc-rich A stars but the number of detections is still low for solar-analogue FGK stars and especially the numerous M stars. For Herschel, we therefore propose the DEBRIS (“Disc Emission via a Bias-free Reconnaissance in the Infrared/Submillimetre”) Key Project which will probe 450 nearby A–M stellar systems for debris and measure Solar System dust levels in debris discs for the first time. The large sample is statistically robust and without bias, providing a rich legacy for debris disc and exo-planet science. This deep, flux-limited survey will obtain PACS 100/160 images of all 450 systems (472 fields due to wide binaries), and it will be possible to resolve discs toward each one with the PACS’ high resolution. DEBRIS includes SPIRE imaging in systems where debris is detected with PACS (a rate of 50% is expected). The key science questions are: (a) which kinds of stars have debris and why?

(b) what are the sizes, temperatures and masses of the debris discs?  
(c) what is the relation of resolved disc structures to the  
exo-planets? and (d) is our Kuiper Belt common or unusual? The  
Herschel DEBRIS Key Project will answer these questions using  
Herschel's high sensitivity, spectral coverage and resolution. Our  
team includes world experts in debris discs and exo-planet science  
from 8 countries.